

Steady states of the spherically symmetric Vlasov-Poisson system as fixed points of a mass-preserving algorithm

Downloaded from: https://research.chalmers.se, 2025-10-14 02:26 UTC

Citation for the original published paper (version of record):

Andreasson, H., Kunze, M., Rein, G. (2026). Steady states of the spherically symmetric Vlasov-Poisson system as fixed points of a mass-preserving algorithm. Nonlinear Analysis: Real World Applications, 88. http://dx.doi.org/10.1016/j.nonrwa.2025.104467

N.B. When citing this work, cite the original published paper.

research.chalmers.se offers the possibility of retrieving research publications produced at Chalmers University of Technology. It covers all kind of research output: articles, dissertations, conference papers, reports etc. since 2004. research.chalmers.se is administrated and maintained by Chalmers Library



Contents lists available at ScienceDirect

Nonlinear Analysis: Real World Applications

journal homepage: www.elsevier.com/locate/nonrwa



Steady states of the spherically symmetric Vlasov-Poisson system as fixed points of a mass-preserving algorithm

Håkan Andréasson alo,*, Markus Kunze, Gerhard Rein c

- ^a Mathematical Sciences, Chalmers University of Technology, University of Gothenburg, SE-41296 Gothenburg, Sweden
- ^b Mathematisches Institut, Universität Köln, Weyertal 86-90, D-50931 Köln, Germany
- ^c Fakultät für Mathematik, Physik und Informatik, Universität Bayreuth, D-95440 Bayreuth, Germany

ARTICLE INFO

Keywords: Vlasov-Poisson system Steady states Einstein-Vlasov system

ABSTRACT

We give a new proof for the existence of spherically symmetric steady states to the Vlasov-Poisson system, following a strategy that has been used successfully to approximate axially symmetric solutions numerically, both to the Vlasov-Poisson system and to the Einstein-Vlasov system. There are several reasons why a mathematical analysis of this numerical scheme is important. A generalization of the present result to the case of flat axially symmetric solutions would prove that the steady states obtained numerically in Andréasson and Rein (2015) do exist. Moreover, in the relativistic case the question whether a steady state can be obtained by this scheme seems to be related to its dynamical stability. This motivates the desire for a deeper understanding of this strategy.

1. Introduction

Spherically symmetric static solutions and axially symmetric stationary solutions of the Vlasov-Poisson system and of the Einstein-Vlasov system have been extensively studied, both analytically and numerically. A key difference exists between the approaches used to obtain solutions in the spherically symmetric case versus the axially symmetric case. In the spherically symmetric setup, the equations can be treated as an initial value problem. Starting with data at r = 0, a nonlinear system of integro-differential equations can be solved directly, at least numerically, allowing for the construction of a wide range of solutions; cf. [1–6]. In contrast, the axially symmetric case requires a different approach. Numerical methods employed in previous works rely on iterative procedures that must converge to produce a solution.

A crucial ingredient for numerical convergence is that the mass of the iterates is kept constant. We call this a mass-preserving algorithm. This algorithm has successfully been used to construct stationary solutions numerically in the axially symmetric case in a number of works, such as [7–12]. An important feature in these works is that the solutions that are obtained are far from being spherically symmetric. This is very different for the axially symmetric solutions whose existence has been proven, since these results rely on the implicit function theorem, perturbing off a spherically symmetric solution [13–16], and hence the resulting solutions are close to being spherically symmetric. In the case of the Vlasov-Poisson system, also flat axially symmetric steady states have been studied. The mass-preserving algorithm was used in [4] to numerically construct steady states with the property that the corresponding rotation curves are flat, in agreement with observations for disk galaxies. Existence results for flat axially symmetric steady states have been obtained in [17,18], but these results do not cover the type of solutions constructed numerically in [4].

The purpose of this work is to analyze the mass-preserving algorithm that has shown to be so powerful for constructing stationary solutions numerically. This algorithm can be formulated as a fixed point problem, and the main achievement of the present paper is

E-mail addresses: hand@chalmers.se (H. Andréasson), mkunze1@uni-koeln.de (M. Kunze), gerhard.rein@uni-bayreuth.de (G. Rein).

https://doi.org/10.1016/j.nonrwa.2025.104467

Received 22 January 2025; Received in revised form 8 July 2025; Accepted 11 July 2025

Available online 26 July 2025

1468-1218/© 2025 The Authors. Published by Elsevier Ltd. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/).

^{*} Corresponding author.

the proof that under suitable assumptions a fixed point to this problem exists in the case of the spherically symmetric Vlasov-Poisson system. We have therefore found a new method to analytically construct steady states.

In view of the discussion above there are more important reasons for the present investigation. We believe that our result can be extended to cover the flat axially symmetric case, which would prove that the solutions obtained numerically in [4] do actually exist. The estimates in the present work are tailored to the spherically symmetric case, and the modifications required for the flat case are not straightforward and pose an interesting open problem. Another desirable generalization of our method concerns the spherically symmetric Einstein-Vlasov system. Although the existence of solutions can be obtained by other means, there is numerical evidence of an intriguing relation between the fact that a steady state can be obtained by the mass-preserving algorithm and its dynamical stability. Indeed, both stable and unstable steady states are studied in [9] and one can try to numerically construct these steady states by applying the mass-preserving algorithm. There is evidence that only in the case of stable steady states the algorithm converges; see also the discussion in [19]. Analytic progress on the mass-preserving algorithm in the relativistic case could shed some light on this observation, and it could even be valuable for making progress on the nonlinear stability problem.

The outline of the paper is as follows. In the next section the fixed point problem for the Vlasov-Poisson system is formulated in quite general terms. In Section 3 the assumption of spherical symmetry is added. In Section 4 a crucial bound on the potential is derived. Finally, in Section 5 the existence of a fixed point is proven and the main result, Theorem 5.2, is established.

2. The setup

Our aim is to prove the existence of spherically symmetric or axially symmetric, possibly flat steady states of the Vlasov-Poisson system by the following general strategy. We prescribe an ansatz for the particle distribution

$$f = \Phi(E, L) \tag{2.1}$$

in terms of the particle energy

$$E = \frac{1}{2}|v|^2 + U(x)$$

and

$$L = |x \times v|^2$$
 or $L = x_1 v_2 - x_2 v_1$,

the modulus squared of angular momentum or its third component, respectively. For a given spatial density $\rho = \rho(x)$, which is spherically or axially symmetric and has finite mass M > 0 and compact support, we define its induced gravitational potential as

$$U_{\rho}(x) = -\int \frac{\rho(y)}{|x - y|} dy.$$

We substitute U_{ρ} into the ansatz (2.1) to obtain a new spatial density

$$\tilde{\rho}(x) = \int \Phi\left(\frac{1}{2}|v|^2 + U_{\rho}(x), L\right) dv.$$

In addition, we define an amplitude

$$A(\tilde{\rho}) = M \left(\int \tilde{\rho}(x) \, dx \right)^{-1},$$

so that the new spatial density $A(\tilde{\rho}) \tilde{\rho}$ again has mass M. Now assume that we have a fixed point ρ^* of the following map, defined on some suitable domain D:

$$T:D\ni\rho\mapsto U_{\rho}\mapsto\tilde{\rho}=\int\varPhi\left(\frac{1}{2}|v|^2+U_{\rho},L\right)\,dv\mapsto A(\tilde{\rho})\,\tilde{\rho}\in D.$$

Then clearly

$$\rho^* = A(\rho^*)^{-1} \int \varPhi\left(\frac{1}{2}|v|^2 + U_{\rho^*}, L\right) \, dv$$

so that $(f^*, \rho^*, U_{\rho^*})$ is a steady state of the Vlasov-Poisson system, where f^* is given by the new ansatz

$$f^* = A(\rho^*)^{-1} \Phi(E, L).$$

The above observation is the basis for several numerical schemes that are used to construct axially symmetric solutions, as described in the introduction. We will show that it can also employed to prove the existence of steady states, at least in the spherically symmetric case.

3. The operator T in the spherically symmetric case

As a proof of concept for the approach described above, we consider spherically symmetric, isotropic steady states, where we restrict ourselves to the polytropic ansatz

$$f(x,v) = (E_0 - E)_{\perp}^{k},$$
 (3.1)

for $E_0 < 0$ denoting some fixed cut-off energy and k > -1, which will be further restricted in what follows. We wish to obtain steady states of fixed, prescribed mass M > 0 supported in a fixed ball of radius R > 0. As explained above, the resulting steady state will actually have a microscopic equation of state where the right hand side in (3.1) is multiplied by some positive constant. To simplify notation we will restrict ourselves to the choices

$$M = 1$$
, $R = 1$, $E_0 = -1$.

Letting $B = B_1(0)$ denote the open ball in \mathbb{R}^3 of radius 1 centered at the origin, we introduce the following set:

$$D := \{ \rho \in L^p(B) \mid \rho \ge 0 \text{ is spherically symmetric and decreasing with}$$

$$\int_{P} \rho \, dx = 1 \}. \tag{3.2}$$

Here p > 3/2, and spherical symmetry means that $\rho(x) = \rho(r)$ where r = |x|; in what follows we will continue to use the identification encoded in this abuse of notation. Moreover, we will think of functions in D to be extended by 0 to the whole space \mathbb{R}^3 . The set D is a closed, convex subset of the Banach space $L^p(B)$. For $\rho \in D$,

$$U_{\rho}(x) = -\int \frac{\rho(y)}{|x - y|} dy = -\frac{4\pi}{r} \int_{0}^{r} \rho(s) s^{2} ds - 4\pi \int_{r}^{\infty} \rho(s) s ds,$$

and as a function of r, U_{ϱ} is increasing. Moreover, $U_{\varrho}(1) = -1$. Hence U_{ϱ} is the unique solution to the boundary value problem

$$\Delta U = 4\pi \rho$$
 on B, $U = -1$ on ∂B .

Since $\rho \in L^p(B)$ it follows that $U_\rho \in W^{2,p}(B)$, and due to p > 3/2 the latter space is compactly embedded into the space $C(\overline{B})$. Hence the mapping $D \ni \rho \mapsto U_\rho \in C(\overline{B})$ is a compact operator.

The function

$$\tilde{\rho}(x) = \int \left(-1 - \frac{1}{2} |v|^2 - U_{\rho}(x) \right)_+^k dv = c_k \left(-1 - U_{\rho}(x) \right)_+^{k+3/2},$$

where

$$c_k := 4\pi\sqrt{2} \int_0^1 \eta^k \sqrt{1-\eta} \, d\eta,$$

is continuous, non-negative, spherically symmetric, and decreasing, and it vanishes outside \overline{B} since $U_{\rho}(x) > -1$ there. Moreover, the mapping $C_b(\overline{B}) \ni U_{\rho} \mapsto \tilde{\rho} \in L^p(B)$ is continuous. If we multiply $\tilde{\rho}$ by

$$A(\tilde{\rho}) = \left(\int \tilde{\rho}(x) dx\right)^{-1} = \left(c_k \int_B \left(-1 - U_{\rho}(x)\right)_+^{k+3/2} dx\right)^{-1}$$

we obtain a density that is again contained in the set D, and we need to control the behavior of this amplitude. This will be done in the next section.

4. A uniform bound close to the center

Lemma 4.1. There exists $r_0 \in]0,1]$ and $\delta > 0$ such that for every $\rho \in D$ the induced potential has the property that $U_{\rho}(r_0) \leq -1 - \delta$. In particular, $U_{\rho}(r) \leq -1 - \delta$ for all $r \in [0,r_0]$.

Proof. For fixed $r_0 \in]0, \frac{1}{10}]$ we assert that

$$\delta = \min \left\{ \frac{1}{10}, \frac{1}{2} \left(\frac{1}{r_*} - 1 \right) \right\} > 0$$

is a suitable choice, where $r_1 \in]r_0, 1[$ satisfies $r_1^3 \ge \frac{4}{5}$. If $4\pi \int_0^{r_0} s^2 \rho(s) ds \ge \frac{11}{10} r_0$, then

$$\begin{split} U_{\rho}(r_0) &= -\frac{4\pi}{r_0} \int_0^{r_0} s^2 \rho(s) \, ds - 4\pi \int_{r_0}^1 s \rho(s) \, ds \\ &\leq -\frac{4\pi}{r_0} \int_0^{r_0} s^2 \rho(s) \, ds \\ &\leq -\frac{11}{10} = -1 - \frac{1}{10} \leq -1 - \delta \end{split}$$

as desired. Thus in what follows we can assume that

$$4\pi \int_0^{r_0} s^2 \rho(s) \, ds < \frac{11}{10} \, r_0. \tag{4.1}$$

Then $1 - r_1^3 \le \frac{1}{5}$ and hence

$$\frac{3}{1-r_1^3} \left(\frac{1}{2} - \frac{11}{10} r_0\right) \frac{r_1^3 - r_0^3}{3} \ge 5 \left(\frac{1}{2} - \frac{11}{100}\right) \left(\frac{4}{5} - \frac{1}{1000}\right) \ge \frac{3}{2}. \tag{4.2}$$

This implies that

$$4\pi \int_{r_0}^{r_1} s^2 \rho(s) \, ds \ge \frac{1}{2}. \tag{4.3}$$

Indeed, suppose that $4\pi \int_{r_0}^{r_1} s^2 \rho(s) ds < \frac{1}{2}$. Then $4\pi \int_0^1 r^2 \rho(r) dr = 1$ together with (4.1) yields

$$\frac{1}{2} < 4\pi \int_0^{r_0} s^2 \rho(s) \, ds + 4\pi \int_{r_1}^1 s^2 \rho(s) \, ds$$
$$< \frac{11}{10} r_0 + \frac{4\pi}{3} \rho(r_1) (1 - r_1^3).$$

In the last estimate we used the fact that $\rho \in D$ by definition of D implies that ρ is decreasing, more precisely, there exists a pointwise defined representative of ρ , which we continue to call ρ , and which has the property that ρ is decreasing on the set $[0,1] \setminus N$ where N is a set of measure zero. By making r_1 slightly bigger if necessary we can make sure that $r_1 \notin N$. Hence

$$\rho(r_1) > \frac{3}{4\pi(1-r_1^3)} \left(\frac{1}{2} - \frac{11}{10}r_0\right).$$

from which we deduce that

$$\rho(r) \ge \frac{3}{4\pi(1-r_1^3)} \left(\frac{1}{2} - \frac{11}{10}r_0\right), \text{ for almost all } r \in [0, r_1].$$

As a consequence, by (4.2),

$$\frac{1}{2} > 4\pi \int_{r_0}^{r_1} s^2 \rho(s) \, ds \ge \frac{3}{1 - r_1^3} \left(\frac{1}{2} - \frac{11}{10} \, r_0 \right) \frac{r_1^3 - r_0^3}{3} \ge \frac{3}{2},$$

and this contradiction completes the proof of (4.3). Let $\alpha = 4\pi \int_{r_0}^{r_1} s^2 \rho(s) \, ds$. Then $\alpha \ge \frac{1}{2}$ by (4.3), and

$$4\pi \int_{r_0}^1 s\rho(s) \, ds = 4\pi \int_{r_0}^{r_1} \frac{s^2}{s} \rho(s) \, ds + 4\pi \int_{r_1}^1 s\rho(s) \, ds$$

$$\geq \frac{4\pi}{r_1} \int_{r_0}^{r_1} s^2 \rho(s) \, ds + 4\pi \int_{r_1}^1 s^2 \rho(s) \, ds$$

$$= \frac{\alpha}{r_1} + 4\pi \int_{r_0}^1 s^2 \rho(s) \, ds - 4\pi \int_{r_0}^{r_1} s^2 \rho(s) \, ds$$

$$= \left(\frac{1}{r_1} - 1\right) \alpha + 4\pi \int_{r_0}^1 s^2 \rho(s) \, ds$$

$$\geq \frac{1}{2} \left(\frac{1}{r_1} - 1\right) + 4\pi \int_{r_0}^1 s^2 \rho(s) \, ds$$

$$\geq \left[\frac{1}{2} \left(\frac{1}{r_1} - 1\right) + 1\right] 4\pi \int_{r_0}^1 s^2 \rho(s) \, ds$$

$$\geq (1 + \delta) 4\pi \int_{r_0}^1 s^2 \rho(s) \, ds.$$

Hence

$$\begin{split} U_{\rho}(r_0) &= -\frac{4\pi}{r_0} \int_0^{r_0} s^2 \rho(s) \, ds - 4\pi \int_{r_0}^1 s \rho(s) \, ds \\ &\leq -\frac{4\pi}{r_0} \int_0^{r_0} s^2 \rho(s) \, ds - (1+\delta) \, 4\pi \int_{r_0}^1 s^2 \rho(s) \, ds \\ &= -\frac{4\pi}{r_0} \int_0^{r_0} s^2 \rho(s) \, ds - (1+\delta) \left(1 - 4\pi \int_0^{r_0} s^2 \rho(s) \, ds\right) \\ &= -(1+\delta) + 4\pi \left(1 + \delta - \frac{1}{r_0}\right) \int_0^{r_0} s^2 \rho(s) \, ds. \end{split}$$

Since $r_0(1+\delta) \leq \frac{1}{10} \frac{11}{10} = \frac{11}{100} \leq 1$, we have $1+\delta-\frac{1}{r_0} \leq 0$, so that $U_\rho(r_0) \leq -1-\delta$, which completes the proof. \square

5. Compactness of T and existence of a fixed point

We continue the analysis of the operator T. Using Lemma 4.1 it follows that

$$\int_{B} \left(-1 - U_{\rho}(x)\right)_{+}^{k+3/2} dx \ge \frac{4\pi}{3} r_{0}^{3} \delta^{k+3/2}.$$

On the other hand.

$$U'_{\rho}(r) = \frac{4\pi}{r^2} \int_0^r \rho(s) \, s^2 ds \le \frac{1}{r^2},$$

and since $U_o(1) = -1$ we deduce that

$$U_{\rho}(r) \ge -\frac{1}{r}$$
.

This implies that

$$\int_{\mathbb{R}} \left(-1 - U_{\rho}(x) \right)_{+}^{k+3/2} dx \le 4\pi \int_{0}^{1} s^{2-k-3/2} ds < \infty,$$

provided that k < 3/2. Thus there exist constants $C_1, C_2 > 0$ such that for every $\rho \in D$ the estimate

$$C_1 \le A(\tilde{p}) \le C_2 \tag{5.1}$$

holds. The mapping $D \ni \rho \mapsto A(\tilde{\rho})$ is continuous, and by the previous estimate it is compact. Together with the results in Section 3 this implies that the operator $T: D \to D$ is compact; it is continuous, and for any bounded sequence $(\rho_n) \subset D$ the sequence $(T(\rho_n))$ has a convergent subsequence.

A further useful observation is as follows.

Lemma 5.1. Let -1 < k < 1/2. Then the range $T(D) \subset L^p(B)$ is bounded.

Proof. Recalling (5.1), we only need to estimate $\|\tilde{\rho}\|_{\rho}$ uniformly on D. For $\rho \in D$ we know that $U_{\rho}(r) \ge -1/r$, and this implies that

$$\|\tilde{\rho}\|_{p}^{p} \le C \int_{0}^{1} r^{2-p(k+3/2)} dr < \infty,$$

provided

$$p < \frac{3}{k + 3/2}.$$

Since we required that p > 3/2, this restricts the possible exponents k of our ansatz to -1 < k < 1/2.

Combining the previous results, we are able to state the main result of this paper.

Theorem 5.2. Let -1 < k < 1/2. Then the operator $T: D \to D$ has a fixed point. In particular, the Vlasov-Poisson system has a steady state for the ansatz function (3.1).

Proof. We have seen above that $D \subset L^p(B)$ is closed and convex, $T: D \to D$ is compact and $T(D) \subset L^p(B)$ is bounded. Let $r = \sup \{ \|T(\rho)\|_p : \rho \in D \}$ and $D_r = D \cap \{\rho \in L^p(B) : \|\rho\|_p \le r \}$. Then $D_r \subset L^p(B)$ is closed, bounded and convex, and $T: D_r \to D_r$ is compact. Hence it has a fixed point by Schauder's Theorem; see [20, Theorem 8.8]. \square

References

- [1] H. Andréasson, On static shells and the Buchdahl inequality for the spherically symmetric Einstein-Vlasov system, Comm. Math. Phys. 274 (2007) 409-425.
- [2] H. Andréasson, D. Fajman, M. Thaller, Models for self-gravitating photon shells and geons, Ann. Henri Poincaré 18 (2017) 681-705.
- [3] H. Andréasson, G. Rein, On the steady states of the spherically symmetric Einstein-Vlasov system, Classical Quantum Gravity 24 (2007) 1809-1832.
- [4] H. Andréasson, G. Rein, On the rotation curves for axially symmetric disk solutions of the Vlasov-Poisson system, Mon. Not. R. Astron. Soc. 446 (2015) 3932–3942.
- [5] T. Ramming, G. Rein, Spherically symmetric equilibria for self-gravitating kinetic or fluid models in the non-relativistic and relativistic case—A simple proof for finite extension, SIAM J. Math. Anal. 45 (2013) 900–914.
- [6] G. Rein, Collisionless kinetic equations from astrophysics—The Vlasov-Poisson system, in: C.M. Dafermos, E. Feireisl (Eds.), Handbook of Differential Equations, in: Evolutionary Equations, vol. 3, Elsevier, 2007.
- [7] E. Ames, H. Andréasson, A. Logg, On axisymmetric and stationary solutions of the self-gravitating Vlasov system, Classical Quantum Gravity 33 (2016) 155008. 26pp. (2016).
- [8] E. Ames, H. Andréasson, A. Logg, Cosmic string and black hole limits of toroidal Vlasov bodies in general relativity, Phys. Rev. D 99 (2019) 024012.
- [9] H. Andréasson, G. Rein, A numerical investigation of the stability of steady states and critical phenomena for the spherically symmetric Einstein-Vlasov system, Classical Quantum Gravity 23 (2006) 3659–3677.
- [10] S.L. Shapiro, S.A. Teukolsky, Equilibrium stellar systems with spindle singularities, Astrophys. J. 388 (1992) 287-300.

- [11] S.L. Shapiro, S.A. Teukolsky, Relativistic stellar systems with spindle singularities, Astrophys. J. 419 (1993) 622-635.
- [12] S.L. Shapiro, S.A. Teukolsky, Relativistic stellar systems with rotation, Astrophys. J. 419 (1993) 636-647.
- [13] H. Andréasson, M. Kunze, G. Rein, Existence of axially symmetric static solutions to the Einstein-Vlasov system, Comm. Math. Phys. 308 (2011) 23-47.
- [14] H. Andréasson, M. Kunze, G. Rein, Rotating, stationary, axially symmetric spacetimes with collisionless matter, Comm. Math. Phys. 329 (2014) 787-808.
- [15] G. Rein, Stationary and static stellar dynamic models with axial symmetry, Nonlinear Anal. 41 (2000) 313-344.
- [16] A. Schulze, Existence of axially symmetric solutions to the Vlasov-Poisson system depending on Jacobi's integral, Commun. Math. Sci. 6 (2008) 711-727.
- [17] R. Firt, G. Rein, Stability of disk-like galaxies. I. Stability via reduction, Analysis 26 (2006) 507-525.
- [18] G. Rein, Flat steady states in stellar dynamics—existence and stability, Comm. Math. Phys. 205 (1999) 229-247.
- [19] E. Ames, H. Andréasson, Stationary solutions of the axially symmetric Einstein-Vlasov system: present status and open problems, Classical Quantum Gravity 41 (2024) 073001.
- [20] K. Deimling, Nonlinear Functional Analysis, Springer, Berlin-New York, 1985.